Notes on the function gsw_enthalpy_diff(SA,CT,p_shallow,p_deep) (which is identical to gsw_enthalpy_diff_CT(SA,CT,p_shallow,p_deep))

Notes written 19th September 2010, updated 3rd April 2011

From appendix A.30 of the TEOS-10 Manual (IOC *et al.* (2010)) we have the following expression for the specific enthalpy of seawater, based on the computationally efficient 25-term expression for specific volume (McDougall *et al.* (2010)),

$$\hat{h}^{48}(S_{A},\Theta,p) = c_{p}^{0}\Theta + 10^{4} \left(\frac{a_{2}}{b_{2}} - \frac{2a_{3}b_{1}}{b_{2}^{2}}\right)p + 10^{4} \frac{a_{3}}{2b_{2}}p^{2}$$

$$+ \frac{M^{*}}{2b_{2}}\ln\left(1 + \frac{2b_{1}}{b_{0}}p + \frac{b_{2}}{b_{0}}p^{2}\right) + \frac{N^{*} - \frac{b_{1}}{b_{2}}M^{*}}{(B-A)}\ln\left(1 + p\frac{b_{2}}{A}\frac{(B-A)}{(B+b_{2}p)}\right).$$
(A.30.6)

This has been written in terms of $A = b_1 - \sqrt{b_1^2 - b_0 b_2}$ and $B = b_1 + \sqrt{b_1^2 - b_0 b_2}$, and the coefficients a_0, a_1, a_2, a_3 and b_0, b_1, b_2 and M^* and N^* which are all defined in McDougall *et al.* (2010) and in the TEOS-10 Manual. All of these coefficients are functions of only Absolute Salinity and Conservative Temperature; that is, they are independent of pressure.

The identical functions $\mathbf{gsw_enthalpy_diff}(SA,CT,p_shallow,p_deep)$ and $\mathbf{gsw_enthalpy_diff_CT}(SA,CT,p_shallow,p_deep)$ return the difference between the specific enthalpy of two seawater parcels, both having the same Absolute Salinity and Conservative Temperature, but having different pressures. The two pressures are labeled p^{de} and p^{sh} (for "deep" and "shallow" respectively) and the $\mathbf{gsw_enthalpy_diff}$ code returns $\hat{h}^{48}\left(S_A,\Theta,p^{\text{de}}\right)-\hat{h}^{48}\left(S_A,\Theta,p^{\text{sh}}\right)$ calculated from Eqn. (A.30.6) according to

$$\begin{split} \hat{h}^{25} \left(S_{\mathrm{A}}, \Theta, p^{\mathrm{de}} \right) - \hat{h}^{25} \left(S_{\mathrm{A}}, \Theta, p^{\mathrm{sh}} \right) &= 10^4 \left(p^{\mathrm{de}} - p^{\mathrm{sh}} \right) \left[\left(\frac{a_2}{b_2} - \frac{2a_3b_1}{b_2^2} \right) + \frac{a_3}{2b_2} \left(p^{\mathrm{de}} + p^{\mathrm{sh}} \right) \right] \\ &+ \frac{M^*}{2b_2} \ln \left[1 + \left(p^{\mathrm{de}} - p^{\mathrm{sh}} \right) \frac{\left[2b_1 + b_2 \left(p^{\mathrm{de}} + p^{\mathrm{sh}} \right) \right]}{\left[b_0 + p^{\mathrm{sh}} \left(2b_1 + b_2 p^{\mathrm{sh}} \right) \right]} \right] \\ &+ \frac{N^* - \frac{b_1}{b_2} M^*}{\left(B - A \right)} \ln \left(1 + \left(p^{\mathrm{de}} - p^{\mathrm{sh}} \right) \frac{b_2 \left(B - A \right)}{\left(B + b_2 p^{\mathrm{de}} \right) \left(A + b_2 p^{\mathrm{sh}} \right)} \right). \end{split}$$

The rms error of the $\hat{\rho}^{48}(S_A,\Theta,p)$ 48-term approximation to the TEOS-10 density over the oceanographic "funnel" is 0.00046 kg m⁻³; this can be compared with the rms uncertainty of 0.004 kg m⁻³ of the underlying laboratory density data to which the TEOS-10 Gibbs function was fitted and to the present uncertainty of at least 0.002 kg m⁻³ in evaluating the effects of non-standard seawater composition on the density of seawater.

References

McDougall T. J., P. M. Barker, R. Feistel and D. R. Jackett, 2011: A computationally efficient 48-term expression for the density of seawater in terms of Conservative Temperature, and related properties of seawater. submitted to *Ocean Science Discussions*.

IOC, SCOR and IAPSO, 2010: *The international thermodynamic equation of seawater* – 2010: *Calculation and use of thermodynamic properties*. Intergovernmental Oceanographic Commission, Manuals and Guides No. 56, UNESCO (English), 196 pp. Available from http://www.TEOS-10.org

Below, for motivation and for reference, is section 3.32 and appendices A.30 and K of the TEOS-10 Manual (IOC *et al.* (2010))

3.32 Pressure to height conversion

When vertically integrating the hydrostatic equation $P_z = -g\rho$ in the context of an ocean model where Absolute Salinity S_A and Conservative Temperature Θ (or potential temperature θ) are piecewise constant in the vertical, the geopotential (Eqn. (3.24.2))

$$\Phi = \Phi^0 - \int_{P_0}^{P} v(p') dP', \qquad (3.32.1)$$

can be evaluated as a series of exact differences. If there are a series of layers of index i separated by pressures p^i and p^{i+1} (with $p^{i+1} > p^i$) then the integral can be expressed (making use of (3.7.5), namely $h_P|_{S_A,\Theta} = \hat{h}_P = v$) as a sum over n layers of the differences in specific enthalpy so that

$$\Phi = \Phi^{0} - \int_{P_{0}}^{P} v(p') dP' = \Phi^{0} - \sum_{i=1}^{n-1} \left[\hat{h}(S_{A}^{i}, \Theta^{i}, p^{i+1}) - \hat{h}(S_{A}^{i}, \Theta^{i}, p^{i}) \right].$$
(3.32.2)

The difference in enthalpy at two different pressures for given values of S_A and Θ is available in the GSW Oceanographic Toolbox via the functions **gsw_enthalpy_diff** (equivalently **gsw_enthalpy_diff_CT**) and **gsw_enthalpy_diff_CT_exact**. The summation of a series of such differences in enthalpy occurs in the GSW functions to evaluate two geostrophic streamfunctions from piecewise-constant vertical property profiles, **gsw_geo_strf_dyn_height_pc** and **gsw_geo_strf_isopycnal_pc**.

The vertical integral of the hydrostatic equation ($P_z = -g \rho$ or $g = -v P_z$) can be written as

$$\int_{0}^{z} g(z') dz = -\int_{P_{0}}^{P} v(p') dP' = -\int_{P_{0}}^{P} \hat{v}(S_{SO}, 0^{\circ}C, p') dP' + \Psi$$

$$= -\hat{h}(S_{SO}, 0^{\circ}C, p) + \Psi,$$
(3.32.3)

where the dynamic height anomaly Ψ is expressed in terms of the specific volume anomaly $\hat{\delta} = \hat{v}(S_A, \Theta, p) - \hat{v}(S_{SO}, 0^{\circ}C, p)$ by

$$\Psi = -\int_{P_0}^{P} \hat{\delta}(p') dP', \qquad (3.32.4)$$

where $P_0 = 101\,325\,\text{Pa}$ is the standard atmosphere pressure. Writing the gravitational acceleration of Eqn. (D.3) as $g = g(\phi, z) = g(\phi, 0)(1 - \gamma z)$, the left-hand side of Eqn.

(3.32.3) becomes $g(\phi,0)(z-\frac{1}{2}\gamma z^2)$, and using the 48-term expression for the specific enthalpy of Standard Seawater, Eqn. (3.32.3) becomes

$$\hat{h}^{48}(S_{SO}, 0^{\circ}C, p) - \Psi + g(\phi, 0)(z - \frac{1}{2}\gamma z^{2}) = 0.$$
 (3.32.5)

This is the equation that is solved for height z in the GSW function $\mathbf{gsw_z_from_p}$. It is traditional to ignore the dynamic height anomaly when converting between pressure and height, and this can be done by simply calling this function with only two arguments, as in $\mathbf{gsw_z_from_p}(p,lat)$. Ignoring Ψ makes a difference to z of up to 3.7m at 5000 dbar. Note that height z is negative in the ocean. When the code is called with three arguments, the third argument is taken to be the dynamic height anomaly Ψ and this is used in the solution of Eqn. (3.32.5). Dynamic height anomaly Ψ can be evaluated using the GSW function $\mathbf{gsw_geo_strf_dyn_height}$.

The GSW function <code>gsw_p_from_z</code> is the exact inverse function of <code>gsw_z_from_p</code>; these functions yield outputs that are consistent with each other to machine precision.

A.30 Computationally efficient 48-term expression for the density of seawater in terms of Θ

Ocean models to date have treated their salinity and temperature variables as being Practical Salinity $S_{\rm P}$ and potential temperature θ . Ocean models that are TEOS-10 compatible need to carry Preformed Salinity S_* and Conservative Temperature Θ as their conservative prognostic variables (as discussed in appendices A.20 and A.21), and they need a computationally efficient expression for density in terms of Absolute Salinity $S_{\rm A}$, Conservative Temperature Θ and pressure p.

Following the work of McDougall *et al.* (2003) and Jackett *et al.* (2006), the TEOS-10 density ρ has been approximated by a 48-term rational. The fitted expression is the ratio of two polynomials of (S_A, Θ, p)

$$\rho \approx \rho^{48} = P_{\text{num}}^{\rho 48} / P_{\text{denom}}^{\rho 48}$$
 (A.30.1)

The density data has been fitted in a "funnel" of data points in (S_A, Θ, p) space which is described in more detail in McDougall *et al.* (2011b). The "funnel" extends to a pressure of 8000 dbar. At the sea surface the "funnel" covers the full range of temperature and salinity while for pressures greater than 6500 dbar, the maximum temperature of the fitted data is 10°C and the minimum Absolute Salinity is 30 g kg⁻¹. That is, the fit has been performed over a region of parameter space which includes water that is approximately 8°C warmer and 5 g kg⁻¹ fresher in the deep ocean than the seawater which exists in the present ocean. Table K.1 of appendix K contains the 48 coefficients of the expression (A.30.1) for density in terms of (S_A, Θ, p) .

As outlined in appendix K, this 48-term rational-function expression for ρ yields the thermal expansion and haline contraction coefficients, α^{Θ} and β^{Θ} , that are essentially as accurate as those derived from the full TEOS-10 Gibbs function for data in the "oceanographic funnel". The sound speed derived by differentiating Eqn. (A.30.1) with respect to pressure has an r.m.s. error in the "funnel" of 0.067 m s⁻¹ whereas TEOS-10 fits the available sound speed data with an rms error of only 0.035 m s⁻¹ (Table O.1 of appendix O), so the sound speed obtained from the 48-term expression for density is not quite as accurate as from the full TEOS-10 expression.

In dynamical oceanography it is the thermal expansion and haline contraction coefficients α^{Θ} and β^{Θ} which are the most important aspects of the equation of state since the "thermal wind" is proportional to $\alpha^{\Theta}\nabla_{\nu}\Theta - \beta^{\Theta}\nabla_{\nu}S_{A}$ and the vertical static stability is given in terms of the buoyancy frequency N by $g^{-1}N^2 = \alpha^{\Theta}\Theta_z - \beta^{\Theta}(S_A)_z$. Hence for dynamical oceanography we may take the 48-term rational function expression for density, Eqn. (A.30.1), as essentially reflecting the full accuracy of TEOS-10. This is confirmed in Fig. A.30.1 where the error in using the 48-term expression for density to calculate the isobaric northward density gradient is shown. The vertical axis on this figure is the magnitude of the difference in the northward isobaric density gradient in the world ocean below 1000m when evaluated using Eqn. (A.30.1) versus using the full TEOS-10 Gibbs function. The scales of the axes of this figure have been chosen to be the same as those of Fig. A.5.1 of appendix A.5 so that the smallness of the errors incurred by using the 48-term density expression can be appreciated. By comparing Figs. A.30.1 and A.5.1 it is clear that the much more important issue is to properly represent the effects of seawater composition on seawater density, and this aspect of ocean science is in its infancy. The rms value of the vertical axis in Fig. A.30.1 is 4.6% of that of Fig. A.5.1.

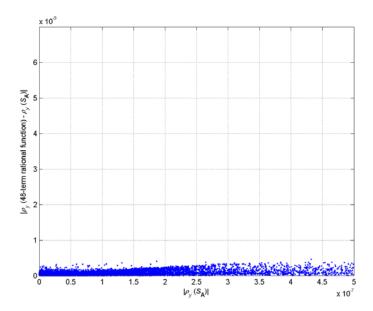


Figure A.30.1. The northward density gradient at constant pressure (the horizontal axis) for data in the world ocean atlas of Gouretski and Koltermann (2004) for p > 1000 dbar. The vertical axis is the magnitude of the difference between evaluating the density gradient using the 48-term expression Eqn. (A.30.1) instead of using the full TEOS-10 expression, using Absolute Salinity $S_{\rm A}$ as the salinity argument in both cases.

Appendix P describes how an expression for the enthalpy of seawater in terms of Conservative Temperature, specifically the functional form $\hat{h}(S_A,\Theta,p)$, together with an expression for entropy in the form $\hat{\eta}(S_A,\Theta)$, can be used as an alternative thermodynamic potential to the Gibbs function $g(S_A,t,p)$. The need for the functional form $\hat{h}(S_A,\Theta,p)$ also arises in section 3.32 and in Eqns. (3.26.3) and (3.29.1). The 48-term expression, Eqn. (A.30.1), for $\rho^{48} = \hat{\rho}^{48}(S_A,\Theta,p)$ can be used to find a closed expression for $\hat{h}(S_A,\Theta,p)$ by integrating the reciprocal of $\hat{\rho}^{48}(S_A,\Theta,p)$ with respect to pressure (in Pa), since $\hat{h}_P = v = \rho^{-1}$ (see Eqn. (2.8.3)).

The 48-term expression for specific volume, Eqn. (A.30.1), is first written explicitly as the ratio of two polynomials in sea pressure p (in dbar) as

$$\hat{v}^{48} = \frac{1}{\hat{\rho}^{48}} = \frac{a_0 + a_1 p + a_2 p^2 + a_3 p^3}{b_0 + 2b_1 p + b_2 p^2},$$
(A.30.2)

where the coefficients a_0 to a_3 and b_0 to b_2 are the following functions of $S_{\rm A}$ and Θ

$$\begin{split} a_0 &= v_{21} + v_{22}\Theta + v_{23}\Theta^2 + v_{24}\Theta^3 + v_{25}\Theta^4 + S_{\rm A} \left(v_{26} + v_{27}\Theta + v_{28}\Theta^2 + v_{29}\Theta^3 + v_{30}\Theta^4\right) \\ &\quad + \left(S_{\rm A}\right)^{1.5} \left(v_{31} + v_{32}\Theta + v_{33}\Theta^2 + v_{34}\Theta^3 + v_{35}\Theta^4\right) + v_{36}S_{\rm A}^2 \;, \\ a_1 &= v_{37} + v_{38}\Theta + v_{39}\Theta^2 + v_{40}\Theta^3 + S_{\rm A} \left(v_{41} + v_{42}\Theta\right), \\ a_2 &= v_{43} + v_{44}\Theta + v_{45}\Theta^2 + v_{46}\Theta S_{\rm A} \;, \\ a_3 &= v_{47} + v_{48}\Theta \;, \\ b_0 &= v_{01} + v_{02}\Theta + v_{03}\Theta^2 + v_{04}\Theta^3 + S_{\rm A} \left(v_{05} + v_{06}\Theta + v_{07}\Theta^2\right) + \left(S_{\rm A}\right)^{1.5} \left(v_{08} + v_{09}\Theta + v_{10}\Theta^2 + v_{11}\Theta^3\right), \\ b_1 &= 0.5 \left(v_{12} + v_{13}\Theta + v_{14}\Theta^2 + S_{\rm A} \left(v_{15} + v_{16}\Theta\right)\right), \\ b_2 &= v_{17} + v_{18}\Theta + v_{19}\Theta^2 + v_{20}S_{\rm A} \;, \end{split}$$

and the numbered coefficients v_1 to v_{48} can be found in Table K.1 (note that $v_{21} = 1$).

It is not difficult to rearrange Eqn. (A.30.2) into the form

$$\hat{v}^{48} = \hat{v}^{48} \left(S_{A}, \Theta, p \right) = \left(\frac{a_2}{b_2} - \frac{2a_3b_1}{b_2^2} \right) + \frac{a_3}{b_2} p + \frac{N + Mp}{b_0 + 2b_1 p + b_2 p^2}, \tag{A.30.3}$$

where N and M are given by

$$N = a_0 + \frac{2a_3b_0b_1}{b_2^2} - \frac{a_2b_0}{b_2}. \quad \text{and} \quad M = a_1 + \frac{4a_3b_1^2}{b_2^2} - \frac{a_3b_0}{b_2} - \frac{2a_2b_1}{b_2}. \quad (A.30.4)$$

The pressure integral of the last term in Eqn. (A.30.3) is well known (see for example section 2.103 of Gradshteyn and Ryzhik (1980)) and is dependent on the sign of the discriminant of the denominator. In our case it can be shown that $b_1^2 > b_0 b_2$ over the full TEOS-10 (S_A, Θ, p) domain, and also that both b_0 is positive while both b_1 and b_2 are negative and bounded away from zero. The indefinite integral, with respect to sea pressure measured in Pa , of the last term in Eqn. (A.30.3) is (with $N^* = 10^4 N$ and $M^* = 10^4 M$)

$$\int \frac{N + Mp}{b_0 + 2b_1 p + b_2 p^2} dP' = \frac{M^*}{2b_2} \ln \left| b_0 + 2b_1 p + b_2 p^2 \right| + \frac{N^* b_2 - M^* b_1}{2b_2 \sqrt{b_1^2 - b_0 b_2}} \ln \left| \frac{b_2 p + b_1 - \sqrt{b_1^2 - b_0 b_2}}{b_2 p + b_1 + \sqrt{b_1^2 - b_0 b_2}} \right|, (A.30.5)$$

The enthalpy $\hat{h}^{48}\left(S_{\rm A},\Theta,p\right)$ is the definite integral of Eqn. (A.30.3) from P_0 to P, plus $c_p^0\Theta$, being the value of enthalpy at P_0 (i. e. at p=0 dbar). Hence the full expression for $\hat{h}^{48}\left(S_{\rm A},\Theta,p\right)$ is (with $A=b_1-\sqrt{b_1^2-b_0b_2}$ and $B=b_1+\sqrt{b_1^2-b_0b_2}$)

$$\hat{h}^{48}(S_{A},\Theta,p) = c_{p}^{0}\Theta + 10^{4} \left(\frac{a_{2}}{b_{2}} - \frac{2a_{3}b_{1}}{b_{2}^{2}}\right)p + 10^{4} \frac{a_{3}}{2b_{2}}p^{2}$$

$$+ \frac{M^{*}}{2b_{2}}\ln\left(1 + \frac{2b_{1}}{b_{0}}p + \frac{b_{2}}{b_{0}}p^{2}\right) + \frac{N^{*} - \frac{b_{1}}{b_{2}}M^{*}}{(B-A)}\ln\left(1 + p\frac{b_{2}}{A}\frac{(B-A)}{(B+b_{2}p)}\right).$$
(A.30.6)

The factor of 10^4 that appears here and in N^* and M^* effectively serves to convert the units of the integration variable from dbar to Pa so that $\hat{h}^{48}(S_A, \Theta, p)$ has units of J kg⁻¹. In these equations S_A is in g kg⁻¹, Θ in °C and p is in dbar. The arguments of the two natural logarithms in Eqn. (A.30.6) are always positive; over the full TEOS-10 (S_A, Θ, p) domain the argument of the first logarithm term is between 0.4 and 1.0 while the argument of the second logarithm term is between 1.0 and 3.5 (note that both b_2 and A are negative while B is positive). Specific enthalpy calculated from Eqn. (A.30.6) is available in the GSW Oceanographic Toolbox as the function gsw_enthalpy(SA,CT,p) (or equivalently **gsw_enthalpy_CT**(SA,CT,p)). The evaluation of $\hat{h}^{48}(S_A,\Theta,p)$ via Eqn. (A.30.6) takes just 12% more computer cpu time than the evaluation of $\hat{v}^{48}(S_A, \Theta, p)$ via a computationally efficient (Hornered in terms of Θ , S_A and p) version of Eqn. (A.30.1). The use of Eqn. (A.30.6) and **gsw_enthalpy** to evaluate $\hat{h}^{48}(S_A, \Theta, p)$ is 9 times faster than first evaluating the in situ temperature t (from $gsw_t_from_CT(SA,CT,p)$) and then calculating enthalpy from the full Gibbs function expression $h(S_A, t, p)$ using gsw_enthalpy_t_exact(SA,t,p). (These last two function calls have also been combined into the one function, **gsw_enthalpy_CT_exact**(SA,CT,p).)

Also, when the enthalpy difference at the same values of S_A and Θ but at different pressures (see Eqn. (3.32.2)) is evaluated using Eqn. (A.30.6), the expression can also be arranged to contain only two logarithm terms (McDougall *et al.* (2011b)). This enthalpy difference is available as the function **gsw_enthalpy_diff**(SA,CT,p) in the GSW Toolbox.

Following Young (2010), the difference between h and $c_p^0\Theta$ may be called "dynamic enthalpy" and can be calculated from Eqn. (A.30.6), recognizing that this equation is based on the 48-term expression for density of McDougall *et al.* (2011b) rather than on the full

TEOS-10 Gibbs function. Dynamic enthalpy is available in the GSW Oceanographic Toolbox as the function $\mathbf{gsw_dynamic_enthalpy}(\mathsf{SA},\mathsf{CT},\mathsf{p})$. Similarly, the partial derivatives of $\hat{h}^{48}(S_\mathsf{A},\Theta,p)$ with respect to Absolute Salinity S_A and with respect to Conservative Temperature Θ can be calculated either by algebraic differentiation of Eqn. (A.30.6) or by first algebraically differentiating Eqn. (A.30.1) and then numerically integrating this expression with respect to pressure (this second procedure is motivated by taking the appropriate S_A or Θ derivatives of Eqn. (3.2.1); see Eqns. (A.18.4) and (A.18.5)) and also Eqns. (A.11.15) and (A.11.18).

Appendix K: Coefficients of 48-term expression for the density of seawater in terms of Θ

The TEOS-10 Gibbs function of seawater $g(S_A, t, p)$ is written as a polynomial in terms of in situ temperature t, while for ocean models, density needs to be expressed as a computationally efficient expression in terms of Conservative Temperature Θ . McDougall $et\ al.$ (2011b) have fitted the TEOS-10 values of density ρ to S_A , Θ and p in a "funnel" of data points in (S_A, Θ, p) space. The fitted expression is in the form of a rational function, being the ratio of two polynomials of (S_A, Θ, p)

$$\rho = P_{\text{num}}^{\rho 48} / P_{\text{denom}}^{\rho 48} . \tag{K.1}$$

The "funnel" of data points in (S_A, Θ, p) space is shown in Figure K.1 and is described in more detail in McDougall *et al.* (2011b); at the sea surface it covers the full range of temperature and salinity while for pressure greater than 6500 dbar, the maximum temperature of the fitted data is 10°C and the minimum Absolute Salinity is 30 g kg⁻¹. The maximum pressure of the "funnel" is 8000 dbar. Table K.1 contains the 48 coefficients of the expression (K.1) for density in terms of (S_A, Θ, p) . The coefficients $v_1 - v_{20}$ in this table have units of kg m⁻³ and the coefficients $v_{21} - v_{48}$ are dimensionless, and the normalizing values of S_A , Θ and p are 1 g kg⁻¹, 1 K and 1 dbar respectively.

The rms error of this 48-term approximation to the TEOS-10 density over the "funnel" is $0.00046~\rm kg~m^{-3}$; this can be compared with the rms uncertainty of $0.004~\rm kg~m^{-3}$ of the underlying laboratory density data to which the TEOS-10 Gibbs function was fitted (see the first two rows of Table O.1 of appendix O). Similarly, the appropriate thermal expansion coefficient,

$$\alpha^{\Theta} = -\frac{1}{\rho} \left. \frac{\partial \rho}{\partial \Theta} \right|_{S_{A}, p}, \tag{K.2}$$

of the 48-term equation of state is different from the same thermal expansion coefficient evaluated from TEOS-10 with an rms error in the "funnel" of $0.069 \, x 10^{-6} \, \text{K}^{-1}$, compared with the rms error of the thermal expansion coefficient of the laboratory data to which the Feistel (2008) Gibbs function was fitted of $0.73 \, x 10^{-6} \, \text{K}^{-1}$ (see row six of Table O.1 of appendix O). In terms of the evaluation of density gradients, the haline contraction coefficient evaluated from Eqn. (K.1) is many times more accurate than the thermal expansion coefficient. Hence we may consider the 48-term rational function expression for density, Eqn. (K.1), to be equally as accurate as the full TEOS-10 expressions for density, for the thermal expansion coefficient and for the saline contraction coefficient for data that reside inside the "oceanographic funnel".

The sound speed evaluated from the 48-term rational function Eqn. (K.1), has an rms error over the "funnel" of $0.067~\rm m~s^{-1}$ which is almost twice the r.m.s. error of the underlying sound speed data that was incorporated into the Feistel (2008) Gibbs function, being $0.035~\rm m~s^{-1}$ (see rows 7 to 9 of Table O.1 of appendix O). Hence, the 48-term expression for density is not quite as accurate as the full TEOS-10 for evaluating sound speed in the ocean. But for dynamical oceanography where α^{Θ} and β^{Θ} are the aspects of the equation of state that, together with spatial gradients of $S_{\rm A}$ and Θ , drive ocean currents and affect the calculation of the buoyancy frequency, we may take the 48-term rational-function expression for density, Eqn. (K.1), as essentially reflecting the full

accuracy of TEOS-10. The accuracy of the 48-term rational function expression for density is illustrated as a function of pressure in Fig. K.2.

The use of Eqn. (K.1) to evaluate $\hat{\rho}(S_A, \Theta, p)$ from $\mathbf{gsw_rho}(SA, CT, p)$ (equivalently $\mathbf{gsw_rho_CT}(SA, CT, p)$) is 6.4 times faster than first evaluating the in situ temperature t (from $\mathbf{gsw_t_from_CT}(SA, CT, p)$) and then calculating in situ density from the full Gibbs function expression $\rho(S_A, t, p)$ via $\mathbf{gsw_rho_t_exact}(SA, t, p)$. (These last two function calls have been combined into $\mathbf{gsw_rho_CT_exact}(SA, CT, P)$.)

	$P_{ m num}^{ ho48}$	Coefficients (kg m ⁻³)		$P_{ m denom}^{ ho48}$	Coefficients (unitless)
v_{01}		9.998 420 897 506 056 x 10 ²	v ₂₁		1.0
v_{02}	Θ	2.839 940 833 161 907 x 10 ⁰	v_{22}	Θ	2.775 927 747 785 646 x 10 ⁻³
v_{03}	Θ^2	-3.147 759 265 588 511 x 10 ⁻²	v_{23}	Θ^2	-2.349 607 444 135 925 x 10 ⁻⁵
v_{04}	Θ^3	1.181 805 545 074 306 x 10 ⁻³	v ₂₄	Θ^3	1.119 513 357 486 743 x 10 ⁻⁶
v_{05}	S_{A}	-6.698 001 071 123 802 x 10 ⁰	v_{25}	Θ^4	6.743 689 325 042 773 x 10 ⁻¹⁰
v_{06}	$S_{\mathrm{A}}\Theta$	-2.986 498 947 203 215 x 10 ⁻²	v_{26}	S_{A}	-7.521 448 093 615 448 x 10 ⁻³
v_{07}	$S_{\rm A}\Theta^2$	2.327 859 407 479 162 x 10 ⁻⁴	v_{27}	$S_{A}\Theta$	-2.764 306 979 894 411 x 10 ⁻⁵
v_{08}	$\left(S_{\rm A}\right)^{1.5}$	-3.988 822 378 968 490 x 10 ⁻²	v_{28}	$S_{\rm A}\Theta^2$	1.262 937 315 098 546 x 10 ⁻⁷
v_{09}	$(S_{\rm A})^{1.5}\Theta$	5.095 422 573 880 500 x 10 ⁻⁴	v_{29}	$S_{\rm A}\Theta^3$	9.527 875 081 696 435 x 10 ⁻¹⁰
v_{10}	$(S_{\rm A})^{1.5} \Theta^2$	-1.426 984 671 633 621 x 10 ⁻⁵	v ₃₀	$S_{\rm A}\Theta^4$	-1.811 147 201 949 891 x 10 ⁻¹¹
<i>v</i> ₁₁	$(S_{\rm A})^{1.5} \Theta^3$	1.645 039 373 682 922 x 10 ⁻⁷	<i>v</i> ₃₁	$\left(S_{\rm A}\right)^{1.5}$	-3.303 308 871 386 421 x 10 ⁻⁵
v_{12}	p	-2.233 269 627 352 527 x 10 ⁻²	v_{32}	$(S_{\rm A})^{1.5}\Theta$	3.801 564 588 876 298 x 10 ⁻⁷
v_{13}	$p\Theta$	-3.436 090 079 851 880 x 10 ⁻⁴	v ₃₃	$\left(S_{\rm A}\right)^{1.5}\Theta^2$	-7.672 876 869 259 043 x 10 ⁻⁹
v_{14}	$p\Theta^2$	3.726 050 720 345 733 x 10 ⁻⁶	v ₃₄	$(S_{\rm A})^{1.5}\Theta^3$	-4.634 182 341 116 144 x 10 ⁻¹¹
<i>v</i> ₁₅	pS_{A}	-1.806 789 763 745 328 x 10 ⁻⁴	v ₃₅	$(S_{\rm A})^{1.5} \Theta^4$	2.681 097 235 569 143 x 10 ⁻¹²
v ₁₆	$p\Theta S_{\mathrm{A}}$	6.876 837 219 536 232 x 10 ⁻⁷	v ₃₆	$S_{\rm A}^2$	5.419 326 551 148 740 x 10 ⁻⁶
v_{17}	p^2	-3.087 032 500 374 211 x 10 ⁻⁷	v ₃₇	p	-2.742 185 394 906 099 x 10 ⁻⁵
v_{18}	$p^2\Theta$	-1.988 366 587 925 593 x 10 ⁻⁸	v ₃₈	$p\Theta$	-3.212 746 477 974 189 x 10 ⁻⁷
v ₁₉	$p^2\Theta^2$	-1.061 519 070 296 458 x 10 ⁻¹¹	v ₃₉	$p\Theta^2$	3.191 413 910 561 627 x 10 ⁻⁹
v_{20}	p^2S_A	1.550 932 729 220 080 x 10 ⁻¹⁰	v_{40}	$p\Theta^3$	-1.931 012 931 541 776 x 10 ⁻¹²
			v_{41}	pS_{A}	-1.105 097 577 149 576 x 10 ⁻⁷
			v_{42}	$p\Theta S_{A}$	6.211 426 728 363 857 x 10 ⁻¹⁰
			v ₄₃	p^2	-1.119 011 592 875 110 x 10 ⁻¹⁰
			v_{44}	$p^2 \Theta$	-1.941 660 213 148 725 x 10 ⁻¹¹
			v ₄₅	$p^2 \Theta^2$	-1.864 826 425 365 600 x 10 ⁻¹⁴
			v ₄₆	$p^2 \Theta S_A$	1.119 522 344 879 478 x 10 ⁻¹⁴
			v_{47}	p^3	-1.200 507 748 551 599 x 10 ⁻¹⁵
			v_{48}	$p^3 \Theta$	6.057 902 487 546 866 x 10 ⁻¹⁷

TABLE K.1 Coefficients of the polynomials $P_{\text{num}}^{\rho 48}\left(S_{\text{A}},\Theta,p\right)$ and $P_{\text{denom}}^{\rho 48}\left(S_{\text{A}},\Theta,p\right)$ that define the 48-term rational-function Eqn. (K.1) for density.

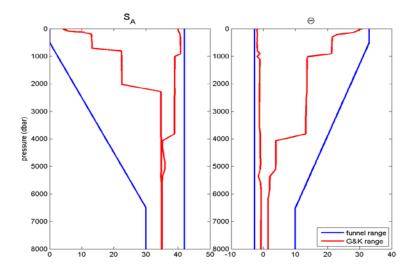


Figure K.1. The ranges of Absolute Salinity and Conservative Temperature in the "Oceanographic funnel" (the blue lines) in which the 48-term expression for density was fitted. The red lines shows the minimum and maximum values of Absolute Salinity and Conservative Temperature that occur in a hydrographic ocean atlas of the world ocean (Gouretski and Koltermann (2004)).

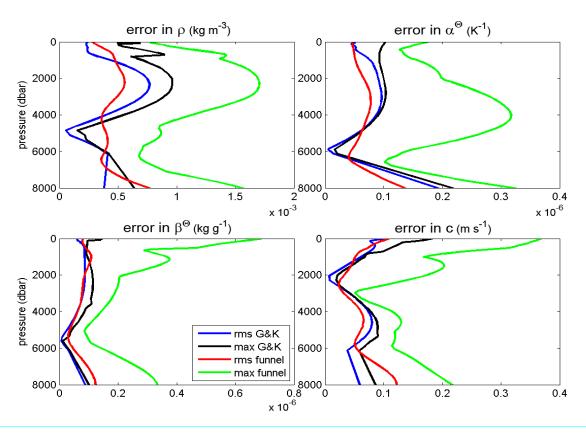


Figure K.2. The errors in using the 48-term rational function expression for density, Eqn. (K.1), to evaluate density, the thermal expansion coefficient, the saline contraction coefficient and sound speed. The red and green lines are the r.m.s. and maximum errors for seawater in the "oceanographic funnel" of McDougall *et al.* (2011b), while the blue and black lines are the r.m.s. and maximum errors for data in the world ocean atlas of Gouretski and Koltermann (2004).